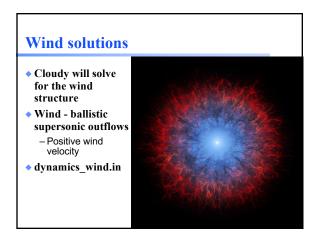
# **Velocity fields**

- Default is static, with thermal broadening
- Turbulence can be added
  - makes line optical depths smaller, so lines escape more easily, continuum florescent excitation more important
- Winds can also be computed
- Line transfer with "Large Velocity Gradient" (LVG) or "Sobolev approximation"
  - $-\,2$  names for same thing



#### Wind solutions

- ~sonic flows from H II regions
  Negative velocity, since motion is towards star
- D-critical flows, nearly at speed of sound
- dynamics\_orion\_flow.in
- Described <u>here</u> and <u>here</u>



### **Project poster**

- One page landscape format PDF with results of the project
- One per group, to be posted on the web site
- Title, authors, abstract
- Introduction
- What problem were you trying to solve?
- Methods and calculations
- Conclusions
- Due by August 18

# **Project poster**

- <u>Non-compliance will be</u> <u>reported to Ted of School</u>
- <u>He has your photo and he</u> <u>knows your address!</u>
- Ted: ted@qub.ac.uk



### Some closing thoughts

- Cloudy a big project, but lots of little projects along the way
- Quantitative spectroscopy read the message in the starlight – what does the spectrum tells us?
- Like all fields, a steep learning curve, but the rewards will be great - be able to decipher the message
  - Like medieval priests, an elevated position since only a few can read the sacred texts