Finding lines in Cloudy

- Run smoke test with command
- **◆** Save line labels
- Spectral label, wavelength, identifies a line
- Save file has label, wavelength, comment about line
- Pick lines from this save file

Luminosity, relative intensity

- Intensity or luminosity of line
 - depending on case
- Intensity relative to normalization line, default $H\boldsymbol{\beta}$
 - Change with normalize command

0	3 88.3323m	-5.577	1.5126
0	3 51.8004m	-5.106	4.4704
0	3 4931.23A	-8.339	0.0026
0	3 4958.91A	-4.876	7.5973
0	3 5006.84A	-4.401	22.6702
0	3 2320.95A	-7.193	0.0366
0	3 4363.21A	-6.593	0.1456
0	3 1660.81A	-7.187	0.0371
0	3 1666.15A	-6.720	0.1087

Emissivity vs density, temperature

• Recombination line, O III forbidden lines

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Two level atom AGN3 Sec 3.5

- Excitation, deexcitation rates
- Transition probabilities
- Critical density
- ◆ Two limits
 - Low densities, every excitation leads to emission of a photon
 - high densities, levels are n LTE, photon emission proportional to $\rm n_{\rm u}~A_{\rm ul}$

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Recombination lines

- ♦ $H^+ + e$ → H^{0*} → $H^0 + photons$
- Critical densities of H I, He I, and He II optical lines are very high, n > 1e15 cm⁻³, so they are usually in LDL
- ◆ Emissivity goes as n²

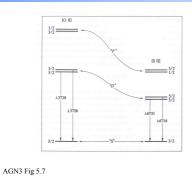
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Forbidden lines

- [O III]
- O⁺⁺ + e → O^{++*} → O⁺⁺ + photons
- ◆ Critical densities of many forbidden lines n ~ 1e3 cm³, so they can be in LDL or HDH
- Emissivity goes as n² or n

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Density indicators



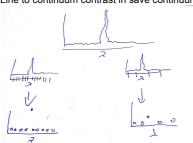
Inward vs total emission

- ◆ "Inwd" label for line
- Inward/outward emission computed on second and later iterations
 - Iterate to convergence
 - Print last

Line to continuum contrast

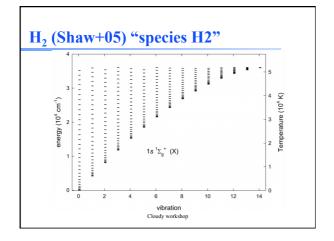
◆ Hazy 1, sec 19.14.44

- Line to continuum contrast in save continuum



Databases in Cloudy

- ◆ Stout (atoms & low ionization)
- Chianti (higher ionization)
- ◆ LAMDA (heavy-element molecules)



Controlling model atoms

- Series of SPECIES XXX commands
- Compare exec time species limit vs small

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